## Matching of correlators in $\mathrm{AdS}_{3} / \mathrm{CFT}_{2}$

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#### Abstract

Recently exact agreement has been found between three-point correlators of (single particle) chiral operators computed in string theory on $A d S_{3} \times S^{3} \times T^{4}$ with NSNS flux and those computed in the symmetric orbifold CFT. However, it has also been shown that these correlators disagree with those computed in supergravity, under any identification of single particle operators which respects the symmetries. In this note we resolve this disagreement: the key point is that mixings with multi-particle operators are not suppressed even at large $N$ in extremal correlators. Allowing for such mixings, orbifold/string theory operators and supergravity operators can be matched such that both non-extremal and extremal three point functions agree, giving further evidence for the non-renormalization of the chiral ring.


Keywords: Gauge-gravity correspondence, AdS-CFT Correspondence.

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## 1. Introduction and summary

Chiral primary operators play an important role in testing the AdS/CFT correspondence. Supersymmetric chiral primaries have protected dimensions, and matching between CFT spectra at weak coupling and supergravity spectra at strong coupling provided the earliest checks of AdS/CFT.

In the correspondence between $\mathcal{N}=4 \mathrm{SYM}$ and string theory on $A d S_{5} \times S^{5}$, the three point functions of $1 / 2$ BPS (single trace) operators computed from supergravity were found to match the corresponding correlators computed in free field theory [1]. This indicated the existence of a previously unknown non-renormalization theorem for such correlators which was subsequently proved, modulo various subtleties, in [2]. Moreover, although four and higher point functions of chiral primaries are in general renormalized, there is evidence that extremal correlators, in which the dimension of one operator is equal to the sum of the others, are also protected [3]; see also the review (4).

It is natural to ask whether similar properties for correlators of chiral primaries hold in the case of $\mathrm{AdS}_{3} / \mathrm{CFT}_{2}$ dualities. The simplest such case is the D1-D5 system, with $n_{1}$ D1-branes and $n_{5}$ D5-branes. Here the duality is between type IIB in an $A d S_{3} \times S^{3} \times X_{4}$ background, where $X_{4}$ is either $T^{4}$ or $K 3$, and a two-dimensional $\mathcal{N}=4$ superconformal field theory; see for example the review (5).

The bulk and boundary theories in this case are known to have equivalent moduli spaces [6], 7], but they are tractable only at distinct points in the moduli space. In the bulk one can work in the supergravity limit, as one does in the case of $\operatorname{AdS} S_{5} \times S^{5}$. One can also consider the S-dual system without RR flux, where the string theory is tractable: for Euclidean $A d S_{3}$ it is described by $H_{3}^{+}$and $\operatorname{SU}(2)$ WZW models at level $k=n_{5}$. The boundary theory becomes tractable in the orbifold limit, namely when the SCFT becomes the symmetric orbifold theory with target space $N=n_{1} n_{5}$ copies of $X_{4}$. Note that the
orbifold theory is not the boundary theory dual to the weakly curved, weakly coupled RR $A d S_{3} \times S^{3} \times X_{4}$ background; the boundary theory is a marginal deformation of the orbifold theory, in which the orbifold is resolved.

Whilst the limits in which the boundary and bulk theories are tractable are at different points in the moduli space, matching of the spectrum of chiral primaries is still possible. Comparison of the spectra obtained from supergravity with those of the boundary theory was first carried out in [8, []. There were also early attempts to compare three point functions computed from supergravity with those computed in the orbifold CFT. Extremal three point functions were computed in the orbifold CFT in [10, 11] whilst the cubic couplings in supergravity relevant for computing three point functions were determined in 12- 14]. It was however noted that these cubic couplings do not match in structure the extremal three point functions computed in the orbifold theory. Only the cubic couplings in supergravity, and not the three point functions, were computed in [13, 14]. Computing the three point functions is rather subtle, in that systematic holographic renormalization (15) is required to obtain the correct correlators, satisfying the requisite Ward identities.

Moreover, extremal correlators are subject to additional subtleties: the bulk extremal cubic couplings vanish, and the corresponding three point functions are obtained from finite boundary terms in the action, which in turn should follow from careful reduction of the ten-dimensional action [3]. Put differently, one should first include boundary terms in the ten-dimensional action such that the variational problem is well-posed for the appropriate Dirichlet boundary conditions, and then dimensionally reduce to obtain the effective threedimensional action.

In practice, it is more convenient to obtain the extremal correlators by analytic continuation of the corresponding non-extremal correlators. That is, one defines the extremal three point functions as

$$
\begin{align*}
\left\langle\mathcal{O}_{\Delta_{2}+\Delta_{3}}\left(x_{1}\right) \mathcal{O}_{\Delta_{2}}\left(x_{2}\right) \mathcal{O}_{\Delta_{3}}\left(x_{3}\right)\right\rangle & =\frac{C_{\left(\Delta_{2}+\Delta_{3}\right) \Delta_{2} \Delta_{3}}}{\left|\overrightarrow{x_{1}}-\overrightarrow{x_{2}}\right|^{2 \Delta_{2}}\left|\overrightarrow{x_{1}}-\overrightarrow{x_{3}}\right|^{2 \Delta_{3}}} \\
C_{\left(\Delta_{2}+\Delta_{3}\right) \Delta_{2} \Delta_{3}} & =\operatorname{Lim}_{\Delta_{1} \rightarrow\left(\Delta_{2}+\Delta_{3}\right)}\left(C_{\Delta_{1} \Delta_{2} \Delta_{3}}\right) \tag{1.1}
\end{align*}
$$

where the scalar operator $\mathcal{O}_{\Delta}$ has dimension $\Delta$, and the non-extremal structure constant $C_{\Delta_{1} \Delta_{2} \Delta_{3}}$ follows from the bulk non-extremal couplings. This analytic continuation was discussed in [3] and more recently such a definition of extremal correlators was discussed in [16] within the framework of holographic renormalization. Note that this approach implicitly assumes that the structure constants are analytic in the operator dimensions, which need not be true, given that the latter are discrete.

Holographically renormalized non-extremal correlators for scalar chiral primaries were recently computed from supergravity in [18], and the corresponding extremal correlators were then determined via analytic continuation. These extremal correlators were compared to those computed in the orbifold CFT. Since only a subset of non-extremal correlators of scalar chiral primaries have so far been computed in the orbifold theory, in 11], only extremal correlators could be compared.

A structural disagreement between these correlators was found. To be more precise, the single particle scalar chiral primaries in the orbifold CFT are labeled by the ( $p, p$ )
cohomology of $X_{4}$, their twist $n \geq 1$, and their R symmetry $\operatorname{SU}(2)_{L} \times \mathrm{SU}(2)_{R}$ quantum numbers as

$$
\begin{equation*}
\mathcal{O}_{n m \bar{m}}^{0,0} ; \quad \mathcal{O}_{n m \bar{m}}^{(r), 1} ; \quad \mathcal{O}_{n m \bar{m}}^{2,2} \tag{1.2}
\end{equation*}
$$

Here $(m, \bar{m})$ are the eigenvalues of $J^{3}$ and $\bar{J}^{3}$ respectively and $(r)$ labels the $(1,1)$ cohomology of $X_{4}$, of dimension $h^{1,1}$; thus $(r)$ runs from 1 to 4 for $T^{4}$ and from 1 to 20 for $K 3$. The operator dimension $\Delta$ is related to the twist and cohomology via

$$
\begin{equation*}
\Delta=(n-1+p), \tag{1.3}
\end{equation*}
$$

with $J=\bar{J}=\frac{1}{2} \Delta$ being the $\mathrm{SO}(4)$ R-symmetry quantum numbers. The cohomology label implicitly defines the transformation properties under the $\mathrm{SO}\left(h^{1,1}\right)$ global symmetry of the CFT.

On the supergravity side one has a set of operators dual to scalar fields in $A d S_{3}$ which are labeled by their dimension $\Delta$ and $R$ symmetry $\mathrm{SU}(2)_{L} \times \mathrm{SU}(2)_{R}$ quantum numbers:

$$
\begin{equation*}
\mathcal{O}_{\Delta m \bar{m}}^{S^{(a)}} ; \quad \mathcal{O}_{\Delta m \bar{m}}^{\Sigma}, \tag{1.4}
\end{equation*}
$$

where $\Delta \geq 1$ for $\mathcal{O}_{\Delta m \bar{m}}^{S^{(a)}}$ and $\Delta \geq 2$ for $\mathcal{O}_{\Delta m \bar{m}}^{\Sigma}$. Here (a) runs from 0 to $h^{1,1}$ and $\Phi \equiv\left(S^{(a)}, \Sigma\right)$ are the bulk scalar fields, which couple to these operators. Of the ( $h^{1,1}+1$ ) operators $\mathcal{O}_{\Delta m \bar{m}}^{S(a)}$ one transforms as a singlet under the $\operatorname{SO}\left(h^{1,1}\right)$ global symmetry and the remaining $h^{1,1}$ as a vector. $\mathcal{O}_{\Delta m \bar{m}}^{\Sigma}$ is also a singlet under the $\mathrm{SO}\left(h^{1,1}\right)$ symmetry; see [9] for further details, and tables of the operators.

Already from (1.2) and (1.4) one can see a subtlety in comparing correlators: the identification between orbifold CFT operators and those dual to supergravity fields is not unique, since the protected quantum numbers of dimension, R symmetry charge and the $\mathrm{SO}\left(h^{1,1}\right)$ global symmetry leave some degeneracy. Given that $\mathcal{O}_{1 m \bar{m}}^{\Sigma}$ is absent, a natural identification between orthonormal operators is

$$
\begin{align*}
& \mathcal{O}_{n m \bar{m}}^{0,0} \leftrightarrow \mathcal{O}_{(n-1) m \bar{m}}^{S} ;  \tag{1.5}\\
& \mathcal{O}_{n m \bar{m}}^{(r) 1,} \leftrightarrow \mathcal{O}_{n m \bar{m}}^{S^{(r)}} ; \\
& \mathcal{O}_{n m \bar{m}}^{2,2} \leftrightarrow \mathcal{O}_{(n+1) m \bar{m}}^{\Sigma},
\end{align*}
$$

and it is this identification which has been assumed in previous literature, but any linear rotation of this identification such that

$$
\begin{equation*}
\binom{\mathcal{O}_{(n-1) m \bar{m}}^{S}}{\mathcal{O}_{(n-1) m \bar{m}}^{\Sigma}}=\mathcal{M}\binom{\mathcal{O}_{n m \bar{m}}^{0,0}}{\mathcal{O}_{(n-2) m \bar{m}}^{2,2}} \tag{1.6}
\end{equation*}
$$

with $\mathcal{M}$ an arbitrary $\mathrm{SO}(2)$ matrix also respects the symmetries and orthonormality.
The result of [18] was that there is a disagreement in the extremal correlators for any choice of $\mathcal{M}$. The disagreement is structural in that for any such linear identification many more of the correlators are non-vanishing in the orbifold theory than in supergravity. Such a disagreement is a priori perhaps not surprising: the computations are at different points in the moduli space and there is no known non-renormalization theorem. Even in
the $\mathcal{N}=4$ SYM theory, which has thirty-two supercharges, the non-renormalization of the analogous three point functions is rather subtle and the proof requires assuming the absence of certain conformal invariants [2]. Renormalization in this case, with only sixteen supercharges, is not a priori excluded, particularly as the orbifold theory is a marginal deformation of the actual boundary CFT.

However, recently three point functions of chiral primaries were computed in the NSNS $A d S_{3} \times S^{3} \times X_{4}$ background using the WZW model description of the worldsheet theory. In [23] extremal three point functions of all single particle chiral primaries were computed, whilst in [22] non-extremal three point functions for operators in the $\mathcal{O}_{n m \bar{m}}^{0,0}$ family were computed. Later in (24 the calculations were extended to non-extremal three point functions of all chiral primaries. All of these correlators agree exactly with those computed in the orbifold CFT, although let us recall that in the latter only a subset of the non-extremal correlators have so far been determined.

That computations at different points in the moduli space agree indicates that there is indeed a non-renormalization theorem protecting these correlators, but at the same time raises the puzzle of why the extremal correlators computed from supergravity did not agree with the orbifold CFT (and hence the string) computations. Whilst it is undoubtedly interesting to find that there is a non-renormalization theorem protecting these correlators, it is arguably more important to understand whether there is any unresolved subtle issue in comparing supergravity and dual field theory results. The reason is that in many situations in gravity/gauge theory dualities one wants to use the supergravity description as a tool to compute the strong coupling result, exactly, when no non-renormalization theorem applies.

In this paper we will resolve this issue, and explain how the supergravity correlators are reconciled with the orbifold CFT and string theory correlators. The conclusions are the following. All non-extremal three point functions computed via supergravity agree precisely with those computed via string theory provided that the matrix $\mathcal{M}$ is such that

$$
\mathcal{M}=\frac{1}{\sqrt{2 \Delta}}\left(\begin{array}{cc}
(\Delta+1)^{1 / 2} & -(\Delta-1)^{1 / 2}  \tag{1.7}\\
(\Delta-1)^{1 / 2} & (\Delta+1)^{1 / 2}
\end{array}\right)
$$

for $\Delta=(n-1) \geq 2$. This agreement provides further evidence for the non-renormalization theorem. Note however that the correspondence between supergravity operators and those in the orbifold CFT is not the naive relation one might have anticipated: $\mathcal{M}$ is not diagonal. This explains the early observation that the cubic couplings in supergravity look very different from the structure constants in the orbifold CFT three point functions.

As discussed in [18] such a linear map between supergravity and orbifold CFT operators is not sufficient to obtain matching for all the extremal correlators. To understand how this issue is resolved, one needs to recall the large $N$ scaling behavior of correlators: the key is that extremal non-linear operator mixings are not suppressed in extremal correlators (3); see also related discussions in Arutyunov:1999en. Thus one can consider an identification between orbifold CFT operators and supergravity operators of the form

$$
\begin{equation*}
\mathcal{O}_{\Delta m \bar{m}}^{p, p} \leftrightarrow \mathcal{O}_{\Delta m \bar{m}}^{\Phi}+\frac{1}{\sqrt{N}} \sum_{i, j} b_{i j} \mathcal{O}_{\Delta_{i} m_{i} \bar{m}_{i}}^{\Phi_{i}} \mathcal{O}_{\left(\Delta-\Delta_{i}\right)\left(m-m_{i}\right)\left(\bar{m}-\bar{m}_{i}\right)}^{\Phi_{j}}+\cdots \tag{1.8}
\end{equation*}
$$

where $b_{i j}$ are certain $N$ independent coefficients and the ellipses denote subleading terms in $N$. Such a two particle term contributes at leading order to certain extremal three point functions, but only at subleading order to non-extremal three point functions. Thus with suitable choices of $b_{i j}$ one can match the extremal correlators computed in supergravity with those computed in string theory and the orbifold theory.

The physical interpretation of such non-linear mixings is that single particle string and orbifold CFT operators do not correspond to single particle supergravity operators. At first sight this may seem surprising, since one might have anticipated that the string worldsheet vertex operators for supergravity modes would correspond to single particle supergravity fields, as they do in flat space. However, there is no contradiction: the matching between supergravity fields and string vertex operators is defined by taking the limit of the string computations of $n$-point functions, and comparing with the corresponding supergravity computations. Thus the comparison made here is the correct way to define the relationship between operators dual to supergravity fields and string vertex operators.

So, to summarize, the non-extremal and extremal three point functions computed in supergravity, in string theory and in the orbifold theory agree provided that one correctly matches operators and takes into account certain extremal non-linear operator mixings. Matching of the correlators determines the map between supergravity and orbifold/string theory operators, where quantum numbers alone do not uniquely determine it.

One might wish to explore whether other correlators are protected by nonrenormalization theorems. In the analogous case of $\mathcal{N}=4 \mathrm{SYM}$, there is evidence that extremal (and next to extremal) n-point functions of chiral primaries are similarly protected, see 4, and thus it is possible that all extremal correlators in this case too will match between string theory, the orbifold CFT and supergravity. Note however that general non-extremal $n$-point functions for $n \geq 4$ are not protected even in $\mathcal{N}=4 \mathrm{SYM}$, and are thus unlikely to be protected in this less supersymmetric system. Comparison of the extremal correlators will again be subtle since non-linear operator mixings may contribute at leading order. That is, in an extremal $n$-point function mixings of the type

$$
\begin{equation*}
\mathcal{O}_{\Delta}^{\Phi}+\frac{1}{N^{(n-2) / 2}} \prod_{i=1}^{n-1} \mathcal{O}_{\Delta_{i}}^{\Phi_{i}}+\cdots \tag{1.9}
\end{equation*}
$$

with $\sum_{i} \Delta_{i}=\Delta$ are not suppressed.
An important open issue is to understand better when there are non-renormalization theorems for correlators. The (almost) proof of the non-renormalization of three point functions in $\mathcal{N}=4$ SYM relies on sophisticated harmonic superspace techniques. In this case one should be able to use the $2 d \mathcal{N}=4$ supersymmetry to explain the nonrenormalization. However, an understanding of the non-renormalization from the bulk supergravity perspective would more immediately generalize to other AdS/CFT dualities. Such a non-renormalization theorem in the bulk would involve arguing that $\alpha^{\prime}$ corrections to the onshell renormalized supergravity action do not contribute to the correlators.

The above discussion relates to three point functions of chiral primaries associated with single particle supergravity fields. Given that these appear to satisfy a non-renormalization
theorem, with appropriate operator mixing taken into account, it seems very likely that all three point functions of multi-particle chiral primaries are similarly protected. In the analogous case of $\mathcal{N}=4 \mathrm{SYM}$, there is indeed evidence for this, from both harmonic superspace considerations [2] and more recently from the holographic analysis of LLM bubbling solutions in [20].

If indeed all three point functions are protected, then an immediate consequence would be that vevs of chiral primary operators in states created by other chiral primaries are also not renormalized. Now in 17-19] such vevs were used to test the proposed correspondence between 1/2 BPS D1-D5 fuzzball geometries and superpositions of RR ground states. Nonrenormalization of the vevs can be used to push this correspondence much further, as will be explored in a separate publication.

The plan of this paper is as follows. In section 2 we review the results of the supergravity computation of three point functions, whilst in section 3 the results of the corresponding string theory computations are reviewed. In section 8 the non-extremal correlators are found to match, with an appropriate linear identification of operators. In section 国 it is shown that additional non-linear terms in this operator identification are needed to obtain matching of extremal correlators.

## 2. Supergravity computation of correlators

In this section we will review the holographic computation of three point functions of single particle scalar chiral primaries. These correlators are computed by perturbing about the $A d S_{3} \times S^{3} \times X_{4}$ background, where $X_{4}$ is $T^{4}$ or $K 3$. There are three distinct families of scalar chiral primaries associated with the $(p, p)$ cohomology of $X_{4}$ with $p=0,1,2$ respectively. The operators couple to the following scalar fields in $A d S_{3}$ :

$$
\begin{equation*}
S_{k I} ; \quad S_{k I}^{(r)} ; \quad \Sigma_{k I} \tag{2.1}
\end{equation*}
$$

Here ( $k, I$ ) denote the $\mathrm{SO}(4) \mathrm{R}$ symmetry labels: $k$ is the degree of the associated $S^{3}$ scalar spherical harmonic and $I$ denotes the remaining Dynkin labels. Expressing $\mathrm{SO}(4)=$ $\mathrm{SU}(2)_{L} \times \mathrm{SU}(2)_{R}$, appropriate labels are the $J_{L / R}^{3}$ charges $(m, \bar{m})$. The middle cohomology of $X_{4}$ is labeled by $(r)=1, \cdots h^{1,1}\left(X_{4}\right)$ where $h^{1,1}$ is four for $T^{4}$ and twenty for $K 3$; this label defines the field transformations under the global $\mathrm{SO}\left(h^{1,1}\right)$ symmetry.

Up to the overall normalization factor, the kinetic terms for these fields are canonically normalized, namely the bulk action is

$$
\begin{align*}
\mathcal{S}=\frac{N}{4 \pi} \int & d^{3} x \sqrt{-G}\left(R_{G}+2-\frac{1}{2}\left(\left(D S_{k I}\right)^{2}-k(k-2)\left(S_{k I}\right)^{2}\right)\right.  \tag{2.2}\\
& \left.-\frac{1}{2}\left(\left(D S_{k I}^{(r)}\right)^{2}-k(k-2)\left(S_{k I}^{(r)}\right)^{2}\right)-\frac{1}{2}\left(\left(D \Sigma_{k I}\right)^{2}-k(k-2) \Sigma_{k I}^{2}\right)+\cdots\right) .
\end{align*}
$$

The overall normalization is proportional to the integer $N=n_{1} n_{5}$. Note that the mass terms are such that the scalar fields associated with degree $k$ harmonics couple to operators of dimension $k$. For the $S$ fields $k \geq 1$ whilst for the $\Sigma$ fields $k \geq 2$. To calculate the three
point functions one also needs the appropriate cubic couplings computed in [12, [13]. These are given by

$$
\begin{align*}
& -\frac{N}{4 \pi} \int d^{3} x \sqrt{-G}\left(T_{123} S^{(a) 1} S^{(a) 2} \Sigma^{3}+U_{123} \Sigma^{1} \Sigma^{2} \Sigma^{3}\right)  \tag{2.3}\\
& \equiv-\frac{N}{16 \pi} \int d^{3} x \sqrt{-G} V_{123} \times \\
& \times\left(\frac{S^{(a) 1} S^{(a) 2} \Sigma^{3}}{\sqrt{\left(k_{1}+1\right)\left(k_{2}+1\right)}}+\frac{\left(k_{1}^{2}+k_{2}^{2}+k_{3}^{2}-2\right)}{\left(k_{1}+1\right)\left(k_{2}+1\right)} \frac{\Sigma^{1} \Sigma^{2} \Sigma^{3}}{6 \sqrt{\left(k_{1}-1\right)\left(k_{2}-1\right)}}\right), \\
& V_{123}=\frac{\Sigma(\Sigma+2)(\Sigma-2) \alpha_{1} \alpha_{2} \alpha_{3} a_{123}}{\left(k_{3}+1\right) \sqrt{k_{1} k_{2} k_{3}\left(k_{3}-1\right)}}
\end{align*}
$$

where $k_{i}$ denotes the dimension of the operator dual to the field $\Psi^{i}, \Sigma=k_{1}+k_{2}+k_{3}$, $\alpha_{1}=\frac{1}{2}\left(k_{2}+k_{3}-k_{1}\right)$ etc and $a_{123}$ is shorthand for the spherical harmonic overlap. Here the label $(a)=1, \cdots h^{1,1}\left(X_{4}\right)+1 \equiv n$ includes all $S$ fields. For subsequent notational convenience we introduce the combinations $T_{123}$ and $U_{123}$ which are defined implicitly by the above equalities. Compactification of type IIB on $X_{4}$ gives rise to a theory with $\mathrm{SO}(n)$ symmetry, and the cubic couplings respect this symmetry. Note however that is an accidental symmetry: only the $\mathrm{SO}\left(h^{(1,1)}\right)$ symmetry is respected by the orbifold CFT and string theory three point functions.

The (renormalized) correlators can be computed using standard holographic renormalization techniques. The two point functions are [18]:

$$
\begin{align*}
\left\langle\mathcal{O}_{k_{1} I_{1}}^{S^{(a)}}(x) \mathcal{O}_{k_{2} I_{2}}^{S^{(b)}}(0)\right\rangle_{h} & =\frac{N}{2 \pi^{2}}\left(k_{1}-1\right)^{2}\left(\frac{1}{x^{2 k_{1}}}\right)_{R} \delta_{I_{1} I_{2}} \delta_{k_{1} k_{2}} \delta^{(a)(b)} ; \quad k \neq 1  \tag{2.4}\\
\left\langle\mathcal{O}_{k_{1} I_{1}}^{\Sigma}(x) \mathcal{O}_{k_{2} I_{2}}^{\Sigma}(0)\right\rangle_{h} & =\frac{N}{2 \pi^{2}}\left(k_{1}-1\right)^{2}\left(\frac{1}{x^{2 k_{1}}}\right)_{R} \delta_{I_{1} I_{2}} \delta_{k_{1} k_{2}}
\end{align*}
$$

where $\mathcal{O}^{S^{(a)}}$ and $\mathcal{O}^{\Sigma}$ denote the operators dual to $S^{(a)}$ and $\Sigma$ respectively. The subscript $R$ indicates that the expressions are renormalized whilst the subscript $h$ in these and subsequent expressions denotes that these are the holographically computed correlators. When $k=1,(k-1)$ is replaced by 1 in the first expression; this is a special case in which the Breitenlohner-Freedman bound is saturated. Recall that there is no $k=1$ operator $\mathcal{O}_{\Sigma}$.

The three point functions are 18]:

$$
\begin{align*}
\left\langle\mathcal{O}^{S^{(a)}}\left(x_{1}\right) \mathcal{O}^{S^{(b)}}\left(x_{2}\right) \mathcal{O}^{\Sigma}\left(x_{3}\right)\right\rangle_{h} & =\frac{N}{4 \pi^{3}} \frac{W_{123} T_{123} \delta^{(a)(b)}}{\left|\vec{x}_{1}-\vec{x}_{2}\right|^{2 \alpha_{3}}\left|\vec{x}_{1}-\vec{x}_{3}\right|^{2 \alpha_{2}}\left|\vec{x}_{2}-\vec{x}_{3}\right|^{2 \alpha_{1}}} ;  \tag{2.5}\\
\left\langle\mathcal{O}^{\Sigma}\left(x_{1}\right) \mathcal{O}^{\Sigma}\left(x_{2}\right) \mathcal{O}^{\Sigma}\left(x_{3}\right)\right\rangle_{h} & =\frac{3 N}{4 \pi^{3}} \frac{W_{123} U_{123}-\left.\vec{x}_{2}\right|^{2 \alpha_{3}}\left|\vec{x}_{1}-\vec{x}_{3}\right|^{2 \alpha_{2}}\left|\vec{x}_{2}-\vec{x}_{3}\right|^{2 \alpha_{1}}}{\mid W_{123}}
\end{align*}=\frac{\Gamma\left(\alpha_{1}\right) \Gamma\left(\alpha_{2}\right) \Gamma\left(\alpha_{3}\right) \Gamma\left(\frac{1}{2}(\Sigma-2)\right)}{\Gamma\left(k_{1}-1\right) \Gamma\left(k_{2}-1\right) \Gamma\left(k_{3}-1\right)} .
$$

Here the operator at position $x_{i}$ has dimension $k_{i}$ and $\mathrm{SO}(4) \mathrm{R}$ symmetry labels $I_{i}$.

To compare with the orbifold CFT and string theory computations one wants normalized three point functions, dividing out by the norms of the operators as given by the two point functions. Suppressing the standard position dependence, this gives:

$$
\begin{align*}
\left\langle\hat{\mathcal{O}}^{S^{(a)}} \hat{\mathcal{O}}^{S^{(b)}} \hat{\mathcal{O}}^{\Sigma}\right\rangle_{h} & =\frac{1}{\sqrt{2 N}} \tilde{W}_{123} T_{123} \delta^{(a)(b)} ;  \tag{2.6}\\
\left\langle\hat{\mathcal{O}}^{\Sigma} \hat{\mathcal{O}}^{\Sigma} \hat{\mathcal{O}}^{\Sigma}\right\rangle_{h} & =\frac{3}{\sqrt{2 N}} \tilde{W}_{123} U_{123} ; \\
\tilde{W}_{123} & =\frac{\Gamma\left(\alpha_{1}\right) \Gamma\left(\alpha_{2}\right) \Gamma\left(\alpha_{3}\right) \Gamma\left(\frac{1}{2}(\Sigma-2)\right)}{\Gamma\left(k_{1}\right) \Gamma\left(k_{2}\right) \Gamma\left(k_{3}\right)},
\end{align*}
$$

where $\hat{O}$ denotes the unit normalized operators. The remaining correlators vanish

$$
\begin{equation*}
\left\langle\hat{\mathcal{O}}^{S^{(a)}} \hat{\mathcal{O}}^{S^{(b)}} \hat{\mathcal{O}}^{S^{(c)}}\right\rangle_{h}=\left\langle\hat{\mathcal{O}}^{S^{(a)}} \hat{\mathcal{O}}^{\Sigma} \hat{\mathcal{O}}^{\Sigma}\right\rangle_{h}=0, \tag{2.7}
\end{equation*}
$$

regardless of the operator dimension.
For later purposes it will be useful to give explicitly extremal correlators in which one of the $\alpha_{i}=0$. These are defined as the continuation of the expressions (2.6): the pole in $\tilde{W}_{123}$ as one of the $\alpha_{i} \rightarrow 0$ cancels a corresponding zero in the bulk couplings ( $T_{123}, U_{123}$ ) to give a finite limit. The relevant normalized extremal three point functions are thus of three types:

$$
\begin{align*}
\left\langle\hat{\mathcal{O}}_{k_{1}+k_{2}}^{S^{(a)} \dagger} \hat{\mathcal{O}}_{k_{1}}^{S^{(b)}} \hat{\mathcal{O}}_{k_{2}}^{\Sigma}\right\rangle_{h} & =\delta^{(a)(b)} \frac{a_{123}}{\sqrt{N}} \sqrt{\frac{2 k_{1} k_{2}\left(k_{1}+k_{2}\right)\left(k_{1}+k_{2}+1\right)}{\left(k_{1}+1\right)\left(k_{2}+1\right)^{2}\left(k_{2}-1\right)}} ;  \tag{2.8}\\
\left\langle\hat{\mathcal{O}}_{k_{1}+k_{2}}^{\Sigma \dagger} \hat{\mathcal{O}}_{k_{1}}^{S(a)} \hat{\mathcal{O}}_{k_{2}}^{S^{(b)}}\right\rangle_{h} & =\delta^{(a)(b)} \frac{a_{123}}{\sqrt{N}} \sqrt{\frac{2 k_{1} k_{2}\left(k_{1}+k_{2}\right)}{\left(k_{1}+1\right)\left(k_{2}+1\right)\left(k_{1}+k_{2}-1\right)}} ; \\
\quad\left\langle\hat{\mathcal{O}}_{k_{1}+k_{2}}^{\Sigma \dagger} \hat{\mathcal{O}}_{k_{1}}^{\Sigma} \hat{\mathcal{O}}_{k_{2}}^{\Sigma}\right\rangle_{h} & =\frac{a_{123}}{\sqrt{N}} \frac{\left(k_{1}^{2}+k_{2}^{2}+\left(k_{1}+k_{2}\right)^{2}-2\right)^{2}}{\left(k_{1}+1\right)\left(k_{2}+1\right)} \sqrt{\frac{k_{1} k_{2}\left(k_{1}+k_{2}\right)}{2\left(k_{1}-1\right)\left(k_{2}-1\right)\left(k_{1}+k_{2}-1\right)}} .
\end{align*}
$$

Note that the triple overlap $a_{123}=1$ when the operator with maximum dimension also has $\mathrm{SO}(2) \mathrm{R}$ charges which are minus the sums of the $\mathrm{SO}(2)$ charges of the other operators. In particular that the extremal correlators at lowest dimension are:

$$
\begin{equation*}
\left\langle\hat{\mathcal{O}}_{2}^{\Sigma \dagger} \hat{\mathcal{O}}_{1}^{S^{(a)}} \hat{\mathcal{O}}_{1}^{S^{(b)}}\right\rangle_{h}=\delta^{(a)(b)} \frac{a_{123}}{\sqrt{N}} ; \quad\left\langle\hat{\mathcal{O}}_{2}^{S^{(c)} \dagger} \hat{\mathcal{O}}_{1}^{S^{(a)}} \hat{\mathcal{O}}_{1}^{S^{(b)}}\right\rangle_{h}=0 . \tag{2.9}
\end{equation*}
$$

## 3. String theory/orbifold CFT correlators

In this section we will review the results for the corresponding correlators computed in string theory and the orbifold CFT. Not all non-extremal correlators have been computed in the orbifold CFT, but those which are known agree with those computed via the string theory, as do all extremal correlators. Here we summarize the results of $[22-24]$ for the string theory computation of three point functions of scalar chiral primaries. Note that general three point functions involving vector chiral primaries associated with the $(0,2)$ and $(2,0)$ cohomology of $X_{4}$ are also given in [23, 24]. We will not consider these here,
since the corresponding holographic correlators have not been computed, but it should be straightforward to extend our discussions to these operators.

The scalar chiral primaries are labeled by the $(p, p)$ cohomology of $X_{4}$, their twist $n \geq 1$, and their R symmetry $\operatorname{SU}(2)_{L} \times \operatorname{SU}(2)_{R}$ quantum numbers as

$$
\begin{equation*}
\mathcal{O}_{n m \bar{m}}^{0,0} ; \quad \mathcal{O}_{n m \bar{m}}^{(r) 1,1} ; \quad \mathcal{O}_{n m \bar{m}}^{2,2} . \tag{3.1}
\end{equation*}
$$

Here $(m, \bar{m})$ are the eigenvalues of $J^{3}$ and $\bar{J}^{3}$ respectively and $(r)$ labels the $(1,1)$ cohomology of $X_{4}$, of dimension $h^{1,1}$; thus ( $r$ ) runs from 1 to 4 for $T^{4}$ and from 1 to 20 for $K 3$. The cohomology label is equivalent to giving the transformation properties under the $\mathrm{SO}\left(h^{1,1}\right)$ global symmetry. The operator dimension is given by

$$
\begin{equation*}
\Delta=(n-1+p), \tag{3.2}
\end{equation*}
$$

with $J=\bar{J}=\frac{1}{2} \Delta$. These operators are orthonormal

$$
\begin{equation*}
\left\langle\mathcal{O}_{n-m-\bar{m}}^{p, p} \mathcal{O}_{n m \bar{m}}^{p, p}\right\rangle_{s}=1, \tag{3.3}
\end{equation*}
$$

with the subscript $s$ denoting that these are string theory correlators. The three point functions can be conveniently expressed as

$$
\begin{align*}
& \left\langle\mathcal{O}_{n_{1} m_{1} \bar{m}_{1}, \bar{m}_{1}}^{\epsilon_{1}} \mathcal{O}_{n_{2} m_{2} \bar{m}_{2}}^{\epsilon_{2}, \bar{\epsilon}_{2}} \mathcal{O}_{n_{3} m_{3} \bar{m}_{3}}^{\epsilon_{3}, \bar{\epsilon}_{3}}\right\rangle_{s}=\frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \frac{\left(\sum_{i=1}^{3} \epsilon_{i} n_{i}+1\right)\left(\sum_{i=1}^{3} \bar{\epsilon}_{i} n_{i}+1\right)}{4\left(n_{1} n_{2} n_{3}\right)^{1 / 2}} \\
& \left\langle\mathcal{O}_{n_{1} m_{1} \bar{m}_{1}}^{(r) 1,1} \mathcal{O}_{n_{2} m_{2} \bar{m}_{2}}^{(s) 1,1} \mathcal{O}_{\left.n_{3} m_{3} \bar{m}_{3}\right\rangle_{s}}^{\epsilon, \bar{\epsilon}}=\frac{1}{\sqrt{N}} \delta^{(r)(s)} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right)\left(\frac{n_{1} n_{2}}{n_{3}}\right)^{1 / 2}\right. \tag{3.4}
\end{align*}
$$

where $\epsilon=(p-1)$ for $p=0,2$. Here

$$
\begin{align*}
L\left(J_{i}, m_{i}\right) & =d_{m_{1}, m_{2}, m_{3}}^{J_{1}, J_{2}, J_{3}} \eta_{J_{i}}\left(\frac{\alpha_{1}!\alpha_{2}!\alpha_{3}!\left(J_{1}+J_{2}+J_{3}+1\right)!}{\left(2 J_{1}\right)!\left(2 J_{2}\right)!\left(2 J_{3}\right)!}\right)^{1 / 2},  \tag{3.5}\\
\eta_{J_{i}} & =(-)^{\frac{1}{2}\left(J_{1}+J_{2}+J_{3}\right)} .
\end{align*}
$$

with $\alpha_{1}=J_{2}+J_{3}-J_{1}=\frac{1}{2}\left(\Delta_{2}+\Delta_{3}-\Delta_{1}\right)$ etc and

$$
d_{m_{1}, m_{2}, m_{3}}^{J_{1}, J_{2}, J_{3}}=\left(\begin{array}{ccc}
J_{1} & J_{2} & J_{3}  \tag{3.6}\\
m_{1} & m_{2} & m_{3}
\end{array}\right)
$$

are the $\mathrm{SU}(2) 3 j$ symbols. Note that $\mathrm{U}(1)$ R-charge conservation enforces that $m_{1}+m_{2}+$ $m_{3}=\bar{m}_{1}+\bar{m}_{2}+\bar{m}_{3}=0$ in the correlators.

## 4. Matching non-extremal correlators

Let us now consider the matching of the correlators (2.6), (2.7) and (3.4). The most general linear identification of operators which respects the symmetries is

$$
\begin{align*}
\mathbb{O}_{n m \bar{m}}^{S^{(r)}} & \leftrightarrow \mathcal{O}_{n m \bar{m}}^{(r) 1,1} ; \\
\binom{\mathbb{O}_{(n-1) m \bar{m}}^{S}}{\mathbb{O}_{(n-1) m \bar{m}}^{\Sigma}} & =\mathcal{M}\binom{\mathcal{O}_{n m \bar{m}}^{0,0}}{\mathcal{O}_{(n-2) m \bar{m}}^{2,2}}, \tag{4.1}
\end{align*}
$$

for an arbitrary $\mathrm{SO}(2)$ matrix $\mathcal{M}$. Here we denote by $\mathbb{O}_{\Delta}^{\Phi}$ operators which for non-extremal correlators are to be identified with the holographic operators $\hat{\mathcal{O}}_{\Delta}^{\Phi}$. For extremal correlators we will need to refine the map between these operators and the holographic operators.

The matrix $\mathcal{M}$ is completely fixed by the vanishing of the correlators

$$
\begin{equation*}
\left\langle\mathbb{O}^{S} \mathbb{O}^{S^{(r)}} \mathbb{O}^{S^{(s)}}\right\rangle=0 \tag{4.2}
\end{equation*}
$$

which implies

$$
\mathcal{M}=\frac{1}{\sqrt{2 \Delta}}\left(\begin{array}{cc}
(\Delta+1)^{1 / 2} & -(\Delta-1)^{1 / 2}  \tag{4.3}\\
(\Delta-1)^{1 / 2} & (\Delta+1)^{1 / 2}
\end{array}\right)
$$

for $\Delta=(n-1) \geq 2$. Clearly there can be no operator mixing at $\Delta=1$, since there are no dimension one $\mathbb{O}^{\Sigma}$ operators; this is hence a special case which will be discussed separately.

Having determined $\mathcal{M}$, there is no further freedom in the operator identification and one can check whether the remaining correlators agree. Forming the appropriate linear combinations of the string theory correlators (3.4), one finds that

$$
\begin{align*}
\left\langle\mathbb{O}^{S^{(a)}} \mathbb{O}^{S^{(b)}} \mathbb{O}^{S^{(c)}}\right\rangle_{s} & =0 ; \quad \Delta_{i} \neq 1,  \tag{4.4}\\
\left\langle\mathbb{O}^{S^{(a)}} \mathbb{O}^{\Sigma} \mathbb{O}^{\Sigma}\right\rangle_{s} & =0 ; \quad \Delta_{1} \neq 1, \\
\left\langle\mathbb{O}^{\Sigma} \mathbb{O}^{S^{(a)}} \mathbb{O}^{S^{(b)}}\right\rangle_{s} & =\frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \delta^{(a)(b)} \frac{\sqrt{2}\left(\Delta_{1} \Delta_{2} \Delta_{3}\right)^{1 / 2}}{\left(\Delta_{1}^{2}-1\right)^{1 / 2}} ;  \tag{4.5}\\
\left\langle\mathbb{O}^{\Sigma} \mathbb{O}^{\Sigma} \mathbb{O}^{\Sigma}\right\rangle_{s} & =\frac{\left(\Delta_{1}^{2}+\Delta_{2}^{2}+\Delta_{3}^{2}-2\right)}{2\left(\left(\Delta_{2}^{2}-1\right)\left(\Delta_{3}^{2}-1\right)\right)^{1 / 2}}\left\langle\mathbb{O}^{\Sigma} \mathbb{O}^{S^{(a)}} \mathbb{O}^{S^{(b)}}\right\rangle_{s} \tag{4.6}
\end{align*}
$$

Here the subscript $s$ denotes that these are linear combinations of the correlators computed in the string theory.

Now let us compare these correlators with the holographic correlators (2.6) and (2.7): the zeroes given in (2.7) are reproduced (except in the special cases involving dimension one operators). Moreover, using (2.6) and noting that

$$
\begin{equation*}
\frac{3 U_{123}}{T_{231}}=\frac{\left(\Delta_{1}^{2}+\Delta_{2}^{2}+\Delta_{3}^{2}-2\right)}{2\left(\left(\Delta_{2}^{2}-1\right)\left(\Delta_{3}^{2}-1\right)\right)^{1 / 2}} \tag{4.7}
\end{equation*}
$$

one sees that the ratio given in (4.6) indeed agrees with that from supergravity. Thus one need only compare the overall normalization of (4.5) with that of the supergravity correlator for all non-extremal correlators to match. Noting that

$$
\begin{equation*}
\tilde{W}_{123} T_{231}=\frac{\alpha_{1}!\alpha_{2}!\alpha_{3}!\left(J_{1}+J_{2}+J_{3}+1\right)!}{\left(2 J_{1}\right)!\left(2 J_{2}\right)!\left(2 J_{3}\right)!} \frac{\sqrt{\Delta_{1} \Delta_{2} \Delta_{3}}}{\left(\Delta_{1}^{2}-1\right)^{1 / 2}} \frac{2 a_{123}}{\left(\left(\Delta_{1}+1\right)\left(\Delta_{2}+1\right)\left(\Delta_{3}+1\right)\right)^{1 / 2}}, \tag{4.8}
\end{equation*}
$$

the holographic correlator (2.6) can be rewritten in terms of the string theory correlator as

$$
\begin{equation*}
\frac{\left\langle\hat{\mathcal{O}}^{\Sigma} \hat{\mathcal{O}}^{S^{(a)}} \hat{\mathcal{O}}^{S^{(b)}}\right\rangle_{h}}{\left\langle\mathbb{O}^{\Sigma} \mathbb{O}^{S^{(a)}} \mathbb{O}^{S^{(b)}}\right\rangle_{s}}=\frac{1}{\eta_{J_{i}} \eta_{\bar{J}_{i}} d_{m_{1}, m_{2}, m_{3}}^{J_{1}, J_{2} J_{3}} d_{\bar{m}_{1}, \bar{m}_{2}, \bar{m}_{3}}^{\bar{J}_{\bar{\prime}}, \bar{J}_{2} \bar{J}_{3}}} \frac{a_{123}}{\left(\left(\Delta_{1}+1\right)\left(\Delta_{2}+1\right)\left(\Delta_{3}+1\right)\right)^{1 / 2}} \tag{4.9}
\end{equation*}
$$

Triple integrals of spherical harmonics can be expressed in terms of $3 j$ symbols; in particular the triple overlap $a_{123}$ can be written as 26]

$$
\begin{equation*}
a_{123}=\eta_{J_{i}} \eta_{\bar{J}_{i}} d_{m_{1}, m_{2}, m_{3}}^{J_{1}, J_{2} J_{3}} d_{\bar{m}_{1}, \bar{m}_{2}, \bar{m}_{3}}^{\bar{J}_{1}, \bar{J}_{2} \bar{J}_{3}}\left(\left(2 J_{1}+1\right)\left(2 J_{2}+1\right)\left(2 J_{3}+1\right)\right)^{1 / 2} \tag{4.10}
\end{equation*}
$$

and thus the normalization of the holographic correlators precisely matches that of the string correlators!

## 5. Matching of exceptional extremal correlators

The linear matching between supergravity and orbifold CFT operators is sufficient for all non-extremal correlators, and most extremal correlators, to match. There remains however a discrepancy for correlators involving dimension one operators, where no linear mixing was possible.

For correlators which involve at least one dimension one operator, some agree with the holographic results, namely

$$
\begin{align*}
\left\langle\mathbb{O}_{\Delta_{2}}^{S} \mathbb{O}_{1}^{S^{(r)}} \mathbb{O}_{\Delta_{3}}^{S^{(s)}}\right\rangle_{s} & =0 ;  \tag{5.1}\\
\left\langle\mathbb{O}_{\Delta_{2}}^{\Sigma} \mathbb{O}_{1}^{S^{(r)}} \mathbb{O}_{\Delta_{3}}^{S_{3}^{(s)}}\right\rangle_{s} & =\frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \delta^{(r)(s)} \frac{\sqrt{2 \Delta_{2} \Delta_{3}}}{\left(\Delta_{2}^{2}-1\right)^{1 / 2}} ; \\
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{\Delta_{2}}^{S} \mathbb{O}_{\Delta_{3}}^{\Sigma}\right\rangle_{s} & =\frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \frac{\sqrt{2 \Delta_{2} \Delta_{3}}}{\left(\Delta_{3}^{2}-1\right)^{1 / 2}},
\end{align*}
$$

but the rest do not:

$$
\begin{align*}
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{1}^{S} \mathbb{O}_{2}^{S}\right\rangle_{s} & =\frac{2}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) ; \\
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{1}^{S} \mathbb{O}_{2}^{\Sigma}\right\rangle_{s} & =\frac{\sqrt{3}}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) ; \\
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{\Delta_{2}}^{S(r)} \mathbb{O}_{\Delta_{3}}^{\left.S_{3}^{(s)}\right\rangle_{s}}=\right. & \frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right)\left(\frac{\Delta_{2} \Delta_{3}}{2}\right)^{1 / 2} ;  \tag{5.2}\\
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{\Delta_{2}}^{S} \mathbb{O}_{\Delta_{3}}^{S}\right\rangle_{s} & =\frac{1}{\sqrt{2 N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \sqrt{\Delta_{2} \Delta_{3}} ; \\
\left\langle\mathbb{O}_{1}^{S} \mathbb{O}_{\Delta_{2}}^{\Sigma} \mathbb{O}_{\Delta_{3}}^{\Sigma}\right\rangle_{s} & =\frac{1}{2 \sqrt{2 N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \frac{\sqrt{\Delta_{2} \Delta_{3}}}{\left(\left(\Delta_{2}^{2}-1\right)\left(\Delta_{3}^{2}-1\right)\right)^{1 / 2}}\left(\Delta_{2}^{2}+\Delta_{3}^{2}-1\right),
\end{align*}
$$

with the corresponding holographic correlators being

$$
\begin{align*}
\left\langle\hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta_{2}}^{S} \hat{\mathcal{O}}_{\Delta_{3}}^{S}\right\rangle_{h} & =\left\langle\hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta_{2}}^{S^{(r)}} \hat{\mathcal{O}}_{\Delta_{3}}^{(s)}\right\rangle_{h}=\left\langle\hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta_{2}}^{\mathcal{O}_{1}} \hat{\mathcal{O}}_{\Delta_{3}}^{\Sigma}\right\rangle_{h}=0  \tag{5.3}\\
\sqrt{N} & \left.\hat{\mathcal{O}}_{2}^{\Sigma}\right\rangle_{h}
\end{align*}=\frac{1}{\sqrt{3 N}} a_{123} \equiv \frac{2}{\sqrt{3}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) .
$$

Note that all these correlators are extremal because the spherical harmonic triple overlaps are only non-zero when $\Delta_{3}=\left(\Delta_{2} \pm 1\right)$. This follows from the addition of $\mathrm{SO}(4)$ representations

$$
\begin{equation*}
\left(\frac{1}{2}, \frac{1}{2}\right) \oplus\left(\frac{\Delta_{2}}{2}, \frac{\Delta_{2}}{2}\right) \rightarrow\left(\frac{\Delta_{2} \pm 1}{2}, \frac{\Delta_{2} \pm 1}{2}\right) . \tag{5.4}
\end{equation*}
$$

Whilst the extremal holographic correlators are (by construction) the analytic continuation of corresponding non-extremal correlators, the string theory correlators in (5.2) are not the analytic continuation of corresponding non-extremal correlators given in (4.5). As we will now explain, this apparent discrepancy between extremal correlators can be resolved by allowing for non-linear operator mixing.

### 5.1 Large $N$ behavior of correlators

Let $\mathcal{O}_{k}^{\Phi}$ denote the operator of dimension $k$ and $\mathrm{SO}(2) \mathrm{R}$ charges $(k / 2, k / 2)$ dual to the supergravity field $\Phi$, where $\Phi=\left(S^{(a)}, \Sigma\right)$. Now denote by

$$
\begin{equation*}
\mathcal{O}_{k}^{[\Phi]_{n}}=\left[\prod_{i=1}^{n} \mathcal{O}_{k_{i}}^{\Phi_{i}}\right] \tag{5.5}
\end{equation*}
$$

the associated protected $n$-particle operators, with dimension $k=\sum_{i} k_{i}$ and R charges $(k / 2, k / 2)$. Here $\left[\mathcal{O}_{k_{i}}^{\Phi_{i}} \cdots\right]$ denotes the highest weight component of the direct product of $\mathrm{SO}(4)$ representations.

The operators $\mathcal{O}_{k}^{[\Phi]_{n}}$ transform in the same $\mathrm{SO}(4)$ representation as the single particle operators $\mathcal{O}_{k}^{\Phi}$ and therefore one would anticipate that there is operator mixing. Although generically operator mixing with multi-particle operators is suppressed in the large $N$ limit, this is not true for operators transforming in the same representations. One can understand this from large $N$ counting arguments as follows.

Consider first correlation functions of single particle operators. The two and three point functions computed from gravity scale as $N$, as given in (2.4) and (2.5), so the normalized two and three point functions scale as one and $1 / \sqrt{N}$ respectively. Four point functions include both disconnected and connected contributions. The former scale as $N^{2}$ and are such that

$$
\begin{equation*}
\left\langle\mathcal{O}_{k_{1}}^{\Phi_{1}}\left(x_{1}\right) \mathcal{O}_{k_{2}}^{\Phi_{2}}\left(x_{2}\right) \mathcal{O}_{k_{3}}^{\Phi_{3}}\left(x_{3}\right) \mathcal{O}_{k_{4}}^{\Phi_{4}}\left(x_{4}\right)\right\rangle=N^{2}\left(\delta^{\Phi_{1} \Phi_{2}} \delta^{\Phi_{3} \Phi_{4}} \frac{\delta\left(k_{1}+k_{2}\right) \delta\left(k_{3}+k_{4}\right)}{x_{12}^{2\left|k_{1}\right|} x_{34}^{2\left|k_{3}\right|}}+\cdots\right), \tag{5.6}
\end{equation*}
$$

where the ellipses denote permutations and numerical factors are suppressed. The scaling as $N^{2}$ follows from the fact that these disconnected contributions are the products of two point functions. Working with unit normalized operators, the disconnected contribution to the four point function thus scales as one. Connected contributions to four point functions however scale as $N$ or, working with unit normalized operators, as $1 / N$. Note that holographic computation of the connected contributions involves both the cubic and quartic couplings, whilst the disconnected contributions follow entirely from the (renormalized) quadratic action.

Now let us consider correlation functions involving multi particle operators. In particular, one can read off the large $N$ behavior of correlators involving double particle operators from the single particle correlators discussed above. The operator product expansion $\mathcal{O}_{k_{1}}^{\Phi_{1}}\left(x_{1}\right) \mathcal{O}_{k_{2}}^{\Phi_{2}}\left(x_{2}\right)$ contains the term

$$
\begin{equation*}
\mathcal{O}_{k_{1}}^{\Phi_{1}}\left(x_{1}\right) \mathcal{O}_{k_{2}}^{\Phi_{2}}\left(x_{2}\right) \rightarrow\left[\mathcal{O}_{k_{1}}^{\Phi_{1}}\left(x_{1}\right) \mathcal{O}_{k_{2}}^{\Phi_{2}}\left(x_{1}\right)\right] \tag{5.7}
\end{equation*}
$$

with unit coefficient as $x_{1} \rightarrow x_{2}$ since the double particle operator is defined by the short distance limit. Thus from the $x_{1} \rightarrow x_{2}$ behavior of correlators one can extract the $N$ scaling of mixed correlators involving both single and multi particle operators. Working
with unit normalized single particle operators this gives

$$
\begin{align*}
\left\langle\hat{\mathcal{O}}_{k}^{\dagger \Phi_{1}}\left(x_{1}\right)\left[\hat{\mathcal{O}}_{k_{1}}^{\Phi_{2}}\left(x_{2}\right) \hat{\mathcal{O}}_{k_{2}}^{\Phi_{3}}\left(x_{2}\right)\right]\right\rangle & \approx \frac{C_{k_{1} k_{2}}^{123}}{\sqrt{N}} ;  \tag{5.8}\\
\left\langle\left[\hat{\mathcal{O}}_{l}^{\Phi_{1}}\left(x_{1}\right) \hat{\mathcal{O}}_{k-l}^{\Phi_{2}}\left(x_{1}\right)\right]^{\dagger} \hat{\mathcal{O}}_{k_{1}}^{\Phi_{3}}\left(x_{2}\right) \hat{\mathcal{O}}_{k_{2}}^{\Phi_{4}}\left(x_{3}\right)\right\rangle & \approx\left(\delta^{l k_{1}} \delta^{\Phi_{1} \Phi_{3}} \delta^{\Phi_{2} \Phi_{4}}+\delta^{l k_{2}} \delta^{\Phi_{1} \Phi_{4}} \delta^{\Phi_{2} \Phi_{3}}\right) ; \\
\left\langle\left[\hat{\mathcal{O}}_{l}^{\Phi_{1}}\left(x_{1}\right) \hat{\mathcal{O}}_{k-l}^{\Phi_{2}}\left(x_{1}\right)\right]^{\dagger} \hat{\mathcal{O}}_{k_{1}}^{\Phi_{3}}\left(x_{2}\right) \hat{\mathcal{O}}_{k_{2}}^{\Phi_{4}}\left(x_{3}\right)\right\rangle & \approx \frac{1}{N} ; \quad l \neq k_{1}, k_{2} ; \\
\left\langle\left[\hat{\mathcal{O}}_{l}^{\Phi_{1}}\left(x_{1}\right) \hat{\mathcal{O}}_{k-l}^{\Phi_{2}}\left(x_{1}\right)\right]^{\dagger}\left[\hat{\mathcal{O}}_{k_{1}}^{\Phi_{3}}\left(x_{2}\right) \hat{\mathcal{O}}_{k_{2}}^{\Phi_{4}}\left(x_{2}\right)\right]\right\rangle & \approx\left(\delta^{l k_{1}} \delta^{\Phi_{1} \Phi_{3}} \delta^{\Phi_{2} \Phi_{4}}+\delta^{l k_{2}} \delta^{\Phi_{1} \Phi_{4}} \delta^{\Phi_{2} \Phi_{3}}\right) ; \\
\left\langle\left[\hat{\mathcal{O}}_{l}^{\Phi_{1}}\left(x_{1}\right) \hat{\mathcal{O}}_{k-l}^{\Phi_{2}}\left(x_{1}\right)\right]^{\dagger}\left[\hat{\mathcal{O}}_{k_{1}}^{\Phi_{3}}\left(x_{2}\right) \hat{\mathcal{O}}_{k_{2}}^{\Phi_{4}}\left(x_{2}\right)\right]\right\rangle & \approx \frac{1}{N} ; \quad l \neq k_{1}, k_{2} .
\end{align*}
$$

where in all cases $k=k_{1}+k_{2}$, and thus the correlators are extremal. The (standard) $x_{i}$ dependence of the correlators is suppressed. Here structure constants $C_{k_{1} k_{2}}^{123}$ follow from the extremal single particle correlators given in (2.8). The second and fourth correlators follow from the disconnected components of the four point functions, whilst the third and fifth correlators pick up contributions only from the connected components and are thus subleading.

This large $N$ counting demonstrates that operator mixings which are extremal are not suppressed in extremal correlators. That is, suppose one considers operators such that

$$
\begin{equation*}
\left(\tilde{\mathcal{O}}_{k_{b}+k_{c}}^{\Phi_{a}}\right)=\hat{\mathcal{O}}_{k_{b}+k_{c}}^{\Phi_{a}}+\frac{1}{\sqrt{N}} b_{k_{b} k_{c}}^{a b c}\left[\hat{\mathcal{O}}_{k_{b}}^{\Phi_{b}} \hat{\mathcal{O}}_{k_{c}}^{\Phi_{c}}\right]+\cdots \tag{5.9}
\end{equation*}
$$

where the ellipses denote three particle and higher mixings. Then by construction

$$
\begin{align*}
\left\langle\left(\tilde{\mathcal{O}}_{k_{a}}^{\Phi_{a}}\right)^{\dagger}\left(\tilde{\mathcal{O}}_{k_{b}}^{\Phi_{b}}\right)\right\rangle & =\delta^{\Phi_{a} \Phi_{b}} \delta_{k_{a} k_{b}}+\mathcal{O}\left(\frac{1}{N}\right) ;  \tag{5.10}\\
\left\langle\left(\tilde{\mathcal{O}}_{k_{b}+k_{c}}^{\Phi_{a}}\right)^{\dagger}\left(\tilde{\mathcal{O}}_{k_{b}}^{\Phi_{b}}\right)\left(\tilde{\mathcal{O}}_{k_{c}}^{\Phi_{c}}\right)\right\rangle & =\frac{1}{\sqrt{N}}\left(C_{k_{b} k_{c}}^{a b c}+b_{k_{b} k_{c}}^{a b c}\right)+\mathcal{O}\left(\frac{1}{N}\right) \equiv \frac{1}{\sqrt{N}} \tilde{C}_{k_{b} k_{c}}^{a b c}+\mathcal{O}\left(\frac{1}{N}\right) .
\end{align*}
$$

Thus to leading order in $N$ the mixed operators have the same two point functions as the single particle operators, and their three point functions still scale as $1 / \sqrt{N}$. However, the structure constants are modified: $C_{k_{b} k_{c}}^{a b c} \rightarrow \tilde{C}_{k_{b} k_{c}}^{a b c}$. Note that the $N$ scaling of the $m$-particle term in the mixing (5.9) is $1 / N^{m / 2}$ such that the $n$-point functions of the mixed operators scale as $N^{(1-n / 2)}$.

In the case at hand, for the exceptional extremal holographic and string correlators to agree, one needs the following quadratic operator mixings:

$$
\begin{align*}
\mathbb{O}_{2}^{\Sigma} & =\hat{\mathcal{O}}_{2}^{\Sigma}+\frac{1}{2 \sqrt{3 N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{1}^{S}+\cdots ;  \tag{5.11}\\
\mathbb{O}_{2}^{S} & =\hat{\mathcal{O}}_{2}^{S}+\frac{1}{\sqrt{N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{1}^{S}+\cdots ; \\
\mathbb{O}_{\Delta+1}^{S} & =\hat{\mathcal{O}}_{\Delta+1}^{S}+\frac{\sqrt{\Delta(\Delta+1)}}{\sqrt{2 N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta}^{S}+\cdots ; \\
\mathbb{O}_{\Delta+1}^{S^{(r)}}= & \hat{\mathcal{O}}_{\Delta+1}^{S^{(r)}}+\frac{\sqrt{\Delta(\Delta+1)}}{\sqrt{2 N}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta}^{S^{(r)}}+\cdots ; \\
\mathbb{O}_{\Delta+1}^{\Sigma} & =\hat{\mathcal{O}}_{\Delta+1}^{\Sigma}+\frac{1}{\sqrt{2 N}} \frac{\Delta(\Delta+1)}{\sqrt{(\Delta-1)(\Delta+2)}} L\left(J_{i}, m_{i}\right) L\left(\bar{J}_{i}, \bar{m}_{i}\right) \hat{\mathcal{O}}_{1}^{S} \hat{\mathcal{O}}_{\Delta}^{\Sigma}+\cdots,
\end{align*}
$$

where in the latter three cases $\Delta \geq 2$. The ellipses denote additional potential mixings, which include both two particle operators involving vector chiral primaries and $n$-particle operators with $n \geq 3$.

Several computations could in principle be used to verify the consistency of these operator identifications. Firstly, one could compute finite $N$ corrections to the supergravity and string theory/orbifold CFT two and three point functions, although on the supergravity side this is currently intractable since only a subset of the requisite corrections to the effective action are known. Secondly, one could compute correlation functions for operators in the same supermultiplets, which are dual to other supergravity fields. These should also be protected, and the operator identifications required for supergravity and string theory/orbifold CFT correlations functions to agree should descend from those given in (5.11).

As previously mentioned, it would be interesting to understand the nonrenormalization better, both from the perspective of the $2 \mathrm{~d} \mathcal{N}=4$ CFT and from supergravity. This could lead to other non-renormalization theorems and give insights into the required operator matching. More generally one would like to explore further the relationship between the supergravity and string theory computations, to understand better the latter. In supergravity there is by now a deep understanding of the holographic renormalization used to remove infinite volume divergences and obtain renormalized correlators. The same volume renormalization is also responsible for the finiteness of correlators in the string computations, but renormalization has not been systematically developed and applied in this context. Moreover, in supergravity there is a natural geometric understanding of the connection between boundary conditions for bulk fields and dual operators, whilst in the string computations the relation proposed in (27) between worldsheet vertex operators and CFT operators is less well understood. Thus insights from the supergravity holographic computations may help to understand further the (successful) hypotheses used in the string computations.

## Acknowledgments

The author would like to thank Ingmar Kanitscheider and Kostas Skenderis for useful discusssions, and the Simons Workshop for hospitality during the completion of this work. The author is supported by NWO, via the Vidi grant "Holography, duality and time dependence in string theory".

## References

[1] S. Lee, S. Minwalla, M. Rangamani and N. Seiberg, Three-point functions of chiral operators in $D=4, N=4$ SYM at large- $N$, Adv. Theor. Math. Phys. 2 (1998) 697 hep-th/9806074.
[2] K.A. Intriligator, Bonus symmetries of $N=4$ super-Yang-Mills correlation functions via AdS duality, Nucl. Phys. B 551 (1999) 575 hep-th/9811047;
K.A. Intriligator and W. Skiba, Bonus symmetry and the operator product expansion of $N=4$ super- Yang-Mills, Nucl. Phys. B 559 (1999) 165 hep-th/9905020; B. Eden, P.S. Howe and P.C. West, Nilpotent invariants in $N=4$ SYM, Phys. Lett. B 463 (1999) 19 hep-th/9905085;
A. Petkou and K. Skenderis, A non-renormalization theorem for conformal anomalies, Nucl. Phys. B 561 (1999) 100 hep-th/9906030;
P.S. Howe, C. Schubert, E. Sokatchev and P.C. West, Explicit construction of nilpotent covariants in $N=4$ SYM, Nucl. Phys. B 571 (2000) 71 hep-th/9910011;
P.J. Heslop and P.S. Howe, OPEs and 3-point correlators of protected operators in $N=4$ SYM, Nucl. Phys. B 626 (2002) 265 hep-th/0107212.
[3] E. D'Hoker, D.Z. Freedman, S.D. Mathur, A. Matusis and L. Rastelli, Extremal correlators in the $A d S / C F T$ correspondence, hep-th/9908160.
[4] E. D'Hoker and D.Z. Freedman, Supersymmetric gauge theories and the AdS/CFT correspondence, hep-th/0201253.
[5] J.R. David, G. Mandal and S.R. Wadia, Microscopic formulation of black holes in string theory, Phys. Rept. 369 (2002) 549 hep-th/0203048.
[6] R. Dijkgraaf, Instanton strings and hyperKähler geometry, Nucl. Phys. B 543 (1999) 545 hep-th/9810210.
[7] F. Larsen and E.J. Martinec, $\mathrm{U}(1)$ charges and moduli in the D1-D5 system, JHEP 06 (1999) 019 hep-th/9905064.
[8] J. de Boer, Six-dimensional supergravity on $S^{3} \times \operatorname{AdS(3)}$ and $2 D$ conformal field theory, Nucl. Phys. B 548 (1999) 139 hep-th/9806104.
[9] S. Deger, A. Kaya, E. Sezgin and P. Sundell, Spectrum of $D=6, N=4 b$ supergravity on $A d S_{3} \times S^{3}$, Nucl. Phys. B 536 (1998) 110 hep-th/9804166.
[10] A. Jevicki, M. Mihailescu and S. Ramgoolam, Gravity from CFT on $S^{N}(X)$ : symmetries and interactions, Nucl. Phys. B 577 (2000) 47 hep-th/9907144.
[11] O. Lunin and S.D. Mathur, Three-point functions for $M(N) / S(N)$ orbifolds with $N=4$ supersymmetry, Commun. Math. Phys. 227 (2002) 385 hep-th/0103169; Correlation functions for $M(N) / S(N)$ orbifolds, Commun. Math. Phys. 219 (2001) 399 hep-th/0006196.
[12] M. Mihailescu, Correlation functions for chiral primaries in $D=6$ supergravity on $A d S_{3} \times S^{3}$, JHEP 02 (2000) 007 hep-th/9910111.
[13] G. Arutyunov, A. Pankiewicz and S. Theisen, Cubic couplings in $D=6 N=4 b$ supergravity on $A d S_{3} \times S^{3}$, Phys. Rev. D 63 (2001) 044024 hep-th/0007061.
[14] A. Pankiewicz, Six-dimensional supergravities and the AdS/CFT correspondence, Diploma Thesis, University of Munich, Germany, October (2000).
[15] S. de Haro, S.N. Solodukhin and K. Skenderis, Holographic reconstruction of spacetime and renormalization in the AdS/CFT correspondence, Commun. Math. Phys. 217 (2001) 595 hep-th/0002230;
M. Bianchi, D.Z. Freedman and K. Skenderis, How to go with an RG flow, JHEP 08 (2001)

041 hep-th/0105276; Holographic renormalization, Nucl. Phys. B 631 (2002) 159 hep-th/0112119;
D.Z. Freedman, S.D. Mathur, A. Matusis and L. Rastelli, Correlation functions in the CFT(d)/AdS (d + 1) correspondence, Nucl. Phys. B 546 (1999) 96 hep-th/9804058;
K. Skenderis, Lecture notes on holographic renormalization, Class. and Quant. Grav. 19 (2002) 5849 hep-th/0209067;
I. Papadimitriou and K. Skenderis, $A d S / C F T$ correspondence and geometry, hep-th/0404176; Correlation functions in holographic RG flows, JHEP 10 (2004) 075 hep-th/0407071.
[16] K. Skenderis and M. Taylor, Kaluza-Klein holography, JHEP 05 (2006) 057 hep-th/0603016.
[17] K. Skenderis and M. Taylor, Fuzzball solutions and D1-D5 microstates, Phys. Rev. Lett. 98 (2007) 071601 hep-th/0609154.
[18] I. Kanitscheider, K. Skenderis and M. Taylor, Holographic anatomy of fuzzballs, JHEP 04 (2007) 023 hep-th/0611171.
[19] I. Kanitscheider, K. Skenderis and M. Taylor, Fuzzballs with internal excitations, JHEP 06 (2007) 056 arXiv:0704.0690.
[20] K. Skenderis and M. Taylor, Anatomy of bubbling solutions, JHEP 09 (2007) 019 arXiv:0706.0216.
[21] D.Z. Freedman, S.D. Mathur, A. Matusis and L. Rastelli, Correlation functions in the $C F T(d) / A d S(d+1)$ correspondence, Nucl. Phys. B 546 (1999) 96 hep-th/9804058.
[22] M.R. Gaberdiel and I. Kirsch, Worldsheet correlators in $A d S_{3} / C F T_{2}$, JHEP 04 (2007) 050 hep-th/0703001.
[23] A. Dabholkar and A. Pakman, Exact chiral ring of $A d S_{3} / C F T_{2}$, hep-th/0703022.
[24] A. Pakman and A. Sever, Exact $N=4$ correlators of $A d S_{3} / C F T_{2}$, Phys. Lett. B 652 (2007) 60 arXiv:0704.3040.
[25] G. Arutyunov and S. Frolov, Some cubic couplings in type IIB supergravity on $A d S_{5} \times S^{5}$ and three-point functions in $S Y M(4)$ at large- $N$, Phys. Rev. D 61 (2000) 064009 hep-th/9907085.
[26] R.E. Cutkosky, Harmonic functions and matrix elements for hyperspherical quantum field models, J. Math. Phys. 25 (1984) 939.
[27] A. Giveon, D. Kutasov and N. Seiberg, Comments on string theory on AdS ${ }_{3}$, Adv. Theor. Math. Phys. 2 (1998) 733 hep-th/9806194;
D. Kutasov, F. Larsen and R.G. Leigh, String theory in magnetic monopole backgrounds, Nucl. Phys. B 550 (1999) 183 hep-th/9812027;
J. de Boer, H. Ooguri, H. Robins and J. Tannenhauser, String theory on $A d S_{3}$, JHEP 12 (1998) 026 hep-th/9812046;
D. Kutasov and N. Seiberg, More comments on string theory on $A d S_{3}$, JHEP 04 (1999) 008 hep-th/9903219;
J.M. Maldacena and H. Ooguri, Strings in $A d S_{3}$ and the $\mathrm{SL}(2, R)$ WZW model. III: correlation functions, Phys. Rev. D 65 (2002) 106006 hep-th/0111180.

